

Lecture 18

1

Sectional, Ricci, Scalar curvature; Einstein Field Equations.

$$\text{Recall } R(X,Y)Z = \nabla_X \nabla_Y Z - \nabla_Y \nabla_X Z - \nabla_{[X,Y]} Z$$

in coordinates $R = \sum R_{ijk}{}^l dx^i \otimes dx^j \otimes dx^k \otimes \frac{\partial}{\partial x^l}$

$$R\left(\frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j}\right) \frac{\partial}{\partial x^k} = \sum_l R_{ijk}{}^l \frac{\partial}{\partial x^l}$$

Sectional curvature: If $v, w \in T_p M$ are orthonormal vectors, the sectional curvature in the plane spanned by v and w is

$$K_p(v, w) := g(R(v, w)w, v) \quad (g(v, w) = 0, \|v\| = \|w\| = 1)$$

If v and w are not assumed orthonormal, the formula is

$$K_p(v, w) := \frac{g(R(v, w)w, v)}{g(v, v)g(w, w) - g(v, w)^2}$$

The symmetries of the curvature tensor imply that the sectional curvature does not depend on the choice of basis, only on the point p and the plane in the tangent space spanned by v, w .

Traces of tensors: If $T: V \rightarrow V$ is a linear transformation, it has a well-defined **trace**: pick a basis $\{e_1, \dots, e_n\}$ of V , let T_i^j be the components so that $T(e_i) = \sum_j T_i^j e_j$, and set

$$\text{tr}(T) := \sum_{i=1}^n T_i^i \quad (\text{sum of "diagonal" components})$$

This does not depend on the choice of basis.

If $g: V \times V \rightarrow \mathbb{R}$ is a nondegenerate inner product, and $B: V \times V \rightarrow \mathbb{R}$ is any bilinear form, then B has a **trace with respect to g** .

$$B(x, y) = g(x, \tilde{B}(y)) \quad \text{for a unique map } \tilde{B}: V \rightarrow V$$

$$\text{define: } \operatorname{tr}_g(B) := \operatorname{tr} \tilde{B}.$$

Pick basis e_1, e_2, \dots, e_n : then $B_{ij} = B(e_i, e_j)$ $g_{ij} = g(e_i, e_j)$

$$\text{and } B_{ij} = g(e_i, \tilde{B}(e_j)) = \sum_k g_{ik} \tilde{B}_j^k$$

$$\text{So } \tilde{B}_j^k = \sum_i g^{ik} B_{ij} \quad (g^{ik} = \text{inverse matrix of } g_{ik})$$

$$\operatorname{tr}_g B = \sum_j \tilde{B}_j^j = \sum_{i,j} g^{ij} B_{ij}$$

Or, if we take e_1, \dots, e_n to be a g -orthonormal basis of V , then

$$\operatorname{tr}_g B = \sum_i B_{ii} = \sum_i B(e_i, e_i)$$

Ricci curvature: Define a bilinear form $\operatorname{Ric}: T_p M \times T_p M \rightarrow \mathbb{R}$

$$\operatorname{Ric}(X, Y) = \operatorname{tr} (V \mapsto R(V, X)Y)$$

Ric is called **Ricci curvature**. The symmetries of R imply that it is a symmetric bilinear form on each tangent space.

Interpretation: $\frac{1}{n-1} \operatorname{Ric}(x, x)$ is the "average" sectional curvature of all two planes containing x .

In coordinates $\text{Ric}\left(\frac{\partial}{\partial x^i}, \frac{\partial}{\partial x^j}\right) = \text{Ric}_{ij} = \sum_{k=1}^n R_{kij}^k$

[I've noticed that this formula does not match standard notation, where $\text{Ric}_{ab} = \sum_c R^c_{acb}$. It's due to a different notation for Riemann curvature.]

Scalar curvature: Define a function $s: M \rightarrow \mathbb{R}$ by

$$s = \text{tr}_g(\text{Ric})$$

so $s(p) = \sum_i \text{Ric}_p(e_i, e_i)$ where e_i is an orthonormal basis of $T_p M$.

$$\text{or } s = \sum_{i,j} g^{ij} \text{Ric}_{ij} = \sum_{i,j,k} g^{ij} R_{kij}^k$$

Roughly, $s(p)$ is an "average" of all sectional curvatures at p .

Importance: • The sectional curvature is essentially equivalent to the full Riemann curvature tensor, in the sense that if we know all sectional curvatures of all two planes at all points, the Riemann curvature tensor can be determined.

- Ricci and scalar curvature carry progressively less information, but they have geometric meaning: they are related to the deviation of volume from the Euclidean case. (cf. Bishop's theorem)
- A very important combination of tensors is the Einstein tensor

$$G = \text{Ric} - \frac{1}{2} s g \quad G_{ij} = \text{Ric}_{ij} - \frac{1}{2} s g_{ij}$$

The fundamental equation of General Relativity is

$$\underbrace{\text{Ric} - \frac{1}{2}sg}_{\text{Einstein tensor}} + \Lambda g = 8\pi T$$

↖ stress-energy tensor

$\Lambda =$ cosmological constant

Here the metric has signature $(-+++)$

- If Λ and T are taken to be zero, then the equation reduces to $\text{Ric} = 0$.
and $n \neq 2$

So the metric is Ricci-flat.

- If $T=0$ but Λ is not, the equations reduce to $\text{Ric} = \lambda g$ (λ constant)

Such metrics are called Einstein metrics, and they are much studied in the positive definite case as well.

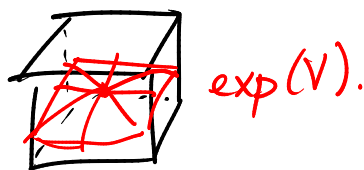
Reason: Take trace of $\text{Ric} - \frac{1}{2}sg + \Lambda g = 0$

$$\text{tr}_g g = n \quad \text{tr}_g \text{Ric} = s \quad \text{so}$$

$$s - \frac{1}{2}sn + \Lambda n = 0 \quad \Leftrightarrow \quad s\left(1 - \frac{n}{2}\right) + \Lambda n = 0$$

$$\Rightarrow s \text{ is constant and } s = \frac{n\Lambda}{\left(\frac{n}{2}-1\right)}$$

Interpretation of sectional curvature: For $v, w \in T_p M$, let $V = \text{span}(v, w) \subset T_p M$
then $K(v, w)$ is the Gaussian curvature at p of the surface $\exp(V) \subset M$



Model Spaces of constant sectional curvature

$$S^n = \{x \in \mathbb{R}^{n+1} \mid \|x\| = 1\}, \quad g = \text{induced metric}, \quad K > 0$$

$$\mathbb{R}^n, \quad g = \text{standard} \quad K = 0$$

$$H^n = \{(x_1, \dots, x_n) \mid x_i > 0\} \quad g = \frac{dx_1^2 + \dots + dx_n^2}{x_n^2}$$

Bishop's Volume Theorem: Suppose (M, g) satisfies $\text{Ric} \geq \rho$ (ρ constant)

(meaning that $\text{Ric}(x, x) \geq \rho$ for every unit vector x .)

Let $E(\rho)$ be the model space of constant curvature whose Ricci curvature equals ρ . (Either S^n, \mathbb{R}^n, H^n with rescaled metric)

Then

$$\text{Vol}(B_M(x, r)) \leq \text{Vol}(B_{E(\rho)}(r))$$

↑
geodesic ball
of radius r
in M

↑
geodesic ball in model
space

Now, if $\rho > 0$, then the model space is S^n , which is compact and has finite volume. Thus if $\text{Ric} \geq \rho > 0$ and (M, g) is complete, M has finite volume.

In fact:

Thm (Bonnet-Myers) if $\text{Ric} \geq \frac{1}{r^2} > 0$ and (M, g) is complete then M is compact and $\text{diam}(M) \leq \pi r$

Cor M has finite fundamental group

Proof The universal cover \tilde{M} also satisfies $\text{Ric} \geq \frac{1}{r^2} > 0$, and so \tilde{M} is compact. Hence the covering map $\tilde{M} \rightarrow M$ can only have finitely many sheets.